





Considerations and applications for an effective high-flux IXS setup

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Main Points For This Talk

Introduce SPring-8 meV-IXS beamlines & some results.

There are significant - even great - opportunities with an "echo" spectrometer.

Some questions & issues worth mentioning.





Operational from ~2002



Baron, et al, J. Phys. Chem. Solids (2000)





BL43LXU Collaborators RIKEN-JASRI Collaboration

Initial Discussions & Design (Beginning in 2004):

Electron Optics: Kouchi SOUTOME, Hitoshi TANAKA Insertion Devices: Takashi TANAKA, Hideo KITAMURA Mono & Cooling: Tetsuro MOCHIZUKI Front End: Sunao TAKAHASHI Hutches and Shielding: Kunikazu TAKESHITA Transport Channel & Optics: Haruhiko OHASHI, Shunji GOTO Spectrometer (2008-): Daisuke ISHIKAWA

More Complete List of Contributors Includes:

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Director/Facilitator: T. ISHIKAWA



SPrin





High Heat-Load Mirror (M1)



Reduces incident beam (<1800W) to <500 W onto LN2 Si(111) Mono

Issues with Silicon Polishing Mirror Mask Design Vacuum Chamber Design Bender Shaft T-Control In-Vac Support T-Control

Note: Vibrations << urad level

Other Major Unexpected Issue

Analyzer Crystals (company quit)





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BL43LXU Analyzer Crystals





9.8 m Radius, 90x94 mm²
50 or 60 μm blade, 3-5 mm depth, ~1 mm pitch Channel width (after etch): ~ 0.15 mm





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RIKEN Temperature Gradient Analyzers

Can remove one of the geometric contributions to the analyzer resolution by making a temperature gradient across the analyzer in the scattering plane.









Res. FWHM [meV]	Energy [keV]	Flux [GHz]	Beam Size <u>VxH</u>	Setup & Comment	
27	15.8	2300	20x45	MR (888) 3 IDs	
>2.8	17.8	100	45x55	HR (999) 3 IDs	
>2.8	17.8	~70	~16x16	HR (999) 3 IDs, µ	
>1.3	21.7	25	45x55	HR (11), 2IDs	
>1.5	21.7	34	45x55	HR (11), 3IDs	
>0.75	25.7	~5	45x55	HR(13) 3 IDs	



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Experiments



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>~1/3 Complex Materials

Usually require a cryostat Sometimes furnace (e.g. ferroelectrics) Often desire to investigate weak modes -> flux limited

<~1/3 Disordered Materials

Resolution (FWHM & tails) generally an issue.

Conditions:

Beam Size: 50-100 um, Standard, ~15 um Possible T: 2K - 800K (specific user groups: 1500K to 3000K) P: 0~200 GPa DACs (User groups: He to 300 Bar) H: up to 7T (BL43) Online area detector for diffraction







Tohoku/RIKEN system









Fukui, et al, JSR 2013

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Tokyo Tech System







Nakajima, Tateno, Imada, Hirose et al





Its Not (only) Carbon In the Core

The earth's is, mostly, but not entirely iron. The density is too low (by ~10%) to be only iron -> some other lighter element must be present.

Based on first HP,HT LIQUID measurements that lighter element can not be simply carbon: the sound velocity would not agree with seismic measurements





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collected at 26 GPa and 2,530 K at momentum transfers Q, as indicated. The spectra include three components: a quasi-elastic peak near zero energy transfer (blue), longitudinal acoustic (LA) phonon mode of liquid $Fe_{84}C_{16}$ (red), and transverse acoustic (TA) phonon mode of diamond (turquoise).

Nakajima et al, Nat Comm 2015 AQRB Sep. 2016









Replace the *figured analyzer* by a *collimating optic* after the sample followed by flat crystal optics. (note also Bortel et al 2000, Sturhahn et al 2011) "Post Sample Collimation"

Shvyd'ko's designs: Refractive Bragg Optics (CDW, CDDW, CDFDW... "echo") Combine Bragg backscattering with refractive angular dispersion. Operate at low (9.1 keV) energy Can have extremely sharp tails on the resolution Parallelization of energy transfers – enhanced efficiency (potentially huge) (Shvyd'ko et al 2006...2016)



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Shvyd'ko's Designs







Spectrometer Applications & Issues

Total Energy Resolution FWHM	Energy Window (Larger is Better)	Q Resolution	Q Range (Larger is Better)	For	Setup
~0.02 meV	>0.2 meV (+) >1 meV (++) >2 meV (+++)	~0.01 nm ⁻¹	$> ~3 \text{ nm}^{-1}$	Glasses, Liquids, QXS Acoustic Phonons & Widths	"Echo"
		~0.10 nm ⁻¹	> 25 nm ⁻¹ (>50)	Optical Phonon Line Widths	
~0.1 meV	>5 meV (+) >10 meV (++) >50 meV (+++)	0.05 - 0.25 nm ⁻¹	$> 25 \text{ nm}^{-1} (>50)$	Great Liquid Instrument	"Echo" (Spectrograph)
		$0.1 - 0.5 \text{ nm}^{-1}$	> 50 nm ⁻¹ (>90)	Great Phonon Instrument	
~0.5 meV	>10 meV (+) > 50 meV (++) > 100 meV (+++)	0.1-0.5 nm ⁻¹	> 50 nm ⁻¹ (>90)	Liquid Survey Instrument	"Echo" & Spectrograph
				Weak Mode Instrument	

 \mathbf{Q}_{max} (9.1 keV, λ =1.36Å) = 92 nm⁻¹ \mathbf{Q} (90 deg) = 65 nm⁻¹ \mathbf{Q} (1 deg) = 0.8 nm⁻¹

Minimum **Q** -> Focus Dependent

















A More Complicated Sample

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Slides with unpublished data removed.

Water and BaBiO3:







Caution



Nominal rate increases are **HUGE**: N or N² where N is 5-500, compared to a conventional spectrometer

But: Energy resolution is much smaller: 1/10 or 1/100 Q resolution may need to be much better: 1/10 or 1/100 Energy is much lower: Losses on thick samples (1/5??) Losses into sample environment (1/5??)

> Very large improvement is possible - but still finite. Care is needed.

(Radiation damage also an issue but dispersive setup helps).





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Absolute Energy Matters (generally higher is better)



Plot shows relative signal for a thicker (10, 100 um) sample. Radiation damage scales similarly (x factor of energy - so slower)



Higher energy worse for Q Resolution





Questions



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What are the resolution function tails like?

What are the limits on sample thickness/projection? Is a grazing incidence geometry on the sample (e.g. thin film) possible?

What to do when you must have a small (1 or 5 or 10 um) spot size? High Pressure, *New Material*

Can one reasonably switch from "echo" to spectrograph operation? For same scan range: just add slit/chan-cut? Or replace mono? Larger scan range possible for spectrograph? Relationship?

Can one gain scan range by increasing the arm length? for both the spectrograph and/or "echo"

Is it possible to (usefully) do this at 20 keV?





Optics Issues

At what level are the designs for different spectrometers (dE, Window, Q, dQ) mutually compatible/exclusive?

Parameter lists including crystal sizes, arm lengths, etc.

Tolerance numbers: Angle/Polishing/Temperature etc.







Two Comments



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The potential of echo/spectrograph operation makes refractive Bragg optics (Shvyd'ko's designs) extremely attractive for both new investigations and significantly improved older style investigations, with the potential to push the state of the art greatly beyond what is possible today.

The spectrometers appear to be difficult practical optics problems, requiring significant funding, time, manpower and experience to make **reliable**.



